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## **Antiproton Production and Energy Density Limitations** in Targets for the Fermilab Pbar Source

I. L. Azhgirey and N. V. Mokhov Institute for High Energy Physics Serpukhov, USSR

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## ANTIPROTON PRODUCTION AND ENERGY DENSITY LIMITATIONS IN TARGETS FOR THE FERMILAB PBAR SOURCE

I.L.AZHGIREY, N.V.MOKHOV

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The recent measurements of the antiproton yield [1] as well as the previous ones [2] differ from the predictions [3] which are the basis of the TEVATRON1 Design Report [4]. It was found in reference [1] that "at small acceptances, where the data depends essentially only on the forward pbar production cross section, the measured yield data indicates that these cross sections were over estimated by about a factor of 3 in the case of tungsten and about 2.3 in the case of copper". To clear up the situation and to understand what one can do to maximize the luminosity of the TEVATRON Collider this work has been done. Two sides of the antiproton production problem are considered:pbar production cross sections and targeting limitations.

Energy deposition density distributions in targets and particle yields are studied via Monte Carlo hadronic and electromagnetic cascade calculations. The calculations of reference [3] rely on the well known CASIM code [5]. In the present work we use two independent Monte Carlo programs:

- MARS10, a fast inclusive program [6] which describes well microscopic and macroscopic data (see ref.[7,8]) in an arbitrary geometry for energies from 10 MeV up to 30 TeV;
- FLUKA86, a program [9] with a full analog simulation of hadronic cascades on the basis of the multichain fragmentation model in the energy region from 50 MeV up to 10 TeV.

Fig.1 shows the antiproton momentum distributions at zero production angle from proton-tungsten collisions at 70 GeV as calculated with the FLUKA86 code and as predicted in refs. [3,10]. Also shown is an experimental data fit from ref[2]. The 80 GeV data of ref[10] have been divided by a factor of 1.2 to take into account an energy dependence of the cross section in this region. FLUKA86 results show no visible difference of the pbar yield in the 0-20 mrad and 0-30 mrad angular intervals for the 6-20 GeV/c momentum region and one can suppose that averaged over these intervals they are compatible with the zero angle data in the considered case. It is confirmed by direct calculation from formula [3] averaged over 0-25 mrad angular interval also shown in graph.

From Fig.1 one can find the following numbers for antiproton yield averaged over the central region (x=0) from 4 to 8 GeV/c:0.05(data[2]), 0.085 (FLUKA86) and 0.10 (formula [3]). The corresponding ratio at 12 Gev/c is 1:2:3.2. It is important that there is no significant Adependence in the FLUKA86 predictions from Al to W, which is in good agreement with the conclusions of ref.[2].

Momentum distributions of antiprotons created in the 0-30 mrad and 0 -60 mrad angular bins when 120 GeV/c protons interact with tungsten nuclei are shown in fig.2. The figure compares curves taken from ref[3] with calculations of the FLUKA86 program. Again the Design Report curves lie higher than the FLUKA86 model predictions.

We have considered the targets and beam parameters which are the same as in the present Fermilab Pbar Source: 5 cm length tungsten and 7 cm length copper targets irradiated by 120 GeV/c protons with 0.1% momentum spread and with 0.7 mrad FWHM divergence. The working point of the beam R.M.S. spot size is 0.378 mm, but for energy density study we considered the 0.1-0.6 mm R.M.S. region.

To prepare a file for the consequent transport calculations in the AP-2 beam line and Debuncher lattice, the produced negatives (p, pi- and K-) from the FLUKA86 program were collected in the various angular intervals with the momentum PO  $\pm$  10% where PO = 8.91058 GeV/c. Table 1 shows the production rates of antiprotons for two targets. The difference in the yields of tungsten and copper targets is less than 20%. The thickness of the targets in units of inelastic scattering lengths (FLUKA) is 0.47 for copper and 0.54 for tungsten. To express the yields in units of pbars per

interacting proton one needs to multiply the W/Cu ratio of yields by factor s = 0.47/0.54 = 0.87. After this the production rates on both targets will be practically the same (see last column of Table 1).

If one runs FLUKA86 at 70 GeV for the conditions of ref[2] one get the following numbers for the antiproton production rate ratios: W/Cu = 1.05 and W/Al=1.11, which coincides with the experiment[2]. The relative yield of other negatives on the Fermilab tungsten and copper targets are given in Table 2 for the 120 mrad angular interval. The results for both materials coincide again.

The practical absence of A-dependence in p-bar yields gives one a chance to examine another target material besides tungsten, more suitable for higher proton beam intensity, to compensate the overestimation of the production cross section in the Design Report.

The smaller proton beam size results in a higher antiproton density in phase space. The maximum energy deposition density  $E_m$  under interactions of the beam with targets, and the physical properties of target itself, determine the minimal spot size. Calculations of three-dimensional energy density distributions in targets intended for antiproton production were performed in ref. [4,11]. One obtained [4] the maximum energy density and tolerable beam intensity for 5 cm tungsten target. We have repeated these calculations with the modern tools FLUKA86 and MARS10 for tungsten and copper targets.

As was mentioned for the first time in ref.[11], when one calculates the energy density distribution for extremely small beam sizes and correspondingly extremely small radial bins special attention should be given to the details of very low energy particle transport which are normally deposited locally. The hadron threshold kinetic energy in CASIM and FLUKA86 is  $E_{\rm th}$ =50 MeV, in MARS10 was  $E_{\rm th}$ =10 MeV. FLUKA86 typically uses an empirical parametrization for electromagnetic showers energy deposition. During the simulation of associated electromagnetic showers in CASIM and MARS10, the electron and photon transport cut-off energy is 100 keV, but the threshold energy at e+e- production and bremstrahlung vertexes was 5 MeV.

Table 3 shows the fraction of the energy deposited locally in each cell. It illustrates the consequence of such approximations for the considered case of antiproton targets: up to 30% of energy in some particular cell could be incorrectly deposited locally. Moreover, in a run with poor statistics, fluctuations in small bins are completely determined by subthreshold particles. We have found the optimum values of the threshold energies: 2 MeV for charge hadron transport and 200 keV for EM-vertexes. We have put them in the present version of MARS10. The result: in all cases the input from locally deposited particles with

energies below the new thresholds is less than 2%.All MARS10 results in this paper have been obtained with this version.

The energies deposited in 5 mm radius and 7 cm length tungsten and copper targets via various chanels calculated with MARS10 and FLUKA86 are given in Table 4. The results of the two programs agree well. The main input comes from electromagnetic showers for both targets: for tungsten number is 85%. Relations between the energy density components in longitudinal distributions differ fairly seriously for the two materials (figs.3,4).

The radial distributions of energy density at the upstream and downstream ends of the tungsten and copper targets are shown in fig.5 and 6. Despite the difference in the threshold energies in MARS10 and FLUKA86, results of these two programs agree very well. For the copper target the results of a special MARS10 run with increased thresholds as well as energy deposition of primary protons at the beginning of the target are given. CASIM data from the Design Report [4], coinciding with others for the first longitudinal bin, lie up to 3 times below MARS10 and FLUKA86 results in the cascade maximum region(fig.5). A. Van Ginneken re-ran expresselly the present version of CASIM and found a much different answer. These new data are shown in Fig.5. Now all three programs are in reasonable agreement.

Usage of the Design Report data resulted in underestimation of the maximum energy density  $E_m$  and correspondingly in an overestimation an admissible beam intensity. Fig.7 shows  $E_m$  as a function of beam R.M.S. spot size. The MARS10 results are fitted well by

$$E_m = C \times \sigma^{-n}$$
, for  $\sigma = \sigma x = \sigma y$  in mm,

where for tungsten C = 0.6 GeV/g, n = 1.0 and for copper C = 0.135 GeV/g, n = 1.35. With the exception of one point the old data [3,4] lie below, but the present CASIM result agrees very well with the MARS10.

There are two possible limits for target behaviour under irradiadiation by high intensity beam: melting point (which is 3410 °C for tungsten and 1083 °C for copper) and shock waves limit. Experts have agreed[12] that for metal targets the last limitation is a priority and that the limit for  $E_{\rm m}$  is 200 Joules/g.Beyond this point shock waves can result in the destruction of the target. The CERN experience confirms this conclusion. The temperature which corresponds to this limit is 1500 °C for tungsten and 760 °C for copper. Now if one knows the maximum energy density  $E_{\rm m}$  in the target per one proton it is easy to

find the tolerable number of beam protons Np. For adiabatic case

$$E_m \times N_D \times 1.6 \times 10^{-10} = 200 \text{ J/g},$$

where  $E_m$  is in GeV/g.

Fig.8 shows  $\,N_{\rm D}\,$  versus beam spot size for the tungsten and copper targets as calculated with MARS10 and fitted by

$${\rm N_p^{=}} \begin{array}{l} 2.1 {\rm x} 10^{12} \ {\rm x} \ \sigma \ {\rm for \ tungsten}, \\ 9.3 {\rm x} 10^{12} \ {\rm x} \ \sigma^{1.35} \ {\rm for \ copper}. \end{array}$$

The Design Report data for tungsten more than twice overestimate our data for  $\sigma > 0.2$  mm, but the present CASIM's point is in excellent agreement again.

So, the copper target should be recomended for the PBAR Source. Tolerable beam intensity on such target is three times higher than for tungsten and antiproton yield is practically the same.

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Table 1. Antiproton production yields from the target.

Collection angle, mrad	Antiprotons, 1.E-	3 per proton Cu	W/Cu	s∗₩/Cu
<120	1.12	.956	1.17	1.02
< 90	1.02	.877	1.16	1.01
< 60	0.81	.673	1.20	1.05
< 30	0.30	.253	1.19	1.03

<sup>\*)</sup> s = 0.47/0.54 = 0.87

Table 2. Relative yield of negatives.

Target	Tungsten	Copper
pbar/pi- pbar/K-	.025	.025
pbar/K-	.30	.28

Table 3. Relative energy deposition at E < Eth.

	:	0 < r < 0.01 cm :					0.01 < r < 0.5 cm			
	:-	P	π±	EM-verte	 x:	P	<b>π</b> ±	EM-vertex		
E <sub>th</sub> , MEV	:-	50	50	5.	:	50	50	5.		
Z, cm	:			TUN	GS1	'EN				
0 - 1	:	.2804	0.	0.	:	.1396	.0100	.0447		
1 - 2	:	.1825	0.	0.	:	.0486	.0137	.1723		
2 - 3	:	.0707	.0100	0.	;	.0316	.0071	.2367		
3 - 4	:	.0949	0.	.3667	:	.0331	.0070	.2180		
4 - 5	:	.0513	.0172	0.	:	.0306	.0095	. 2276		
5 - 6	:	.1179	.0154	0.	:	.0283	.0055	. 2787		
6 - 7	:	.0591	0.	0.	:	.0332	.0219	. 2257		
Z, cm	:			C	OPI	PER				
0 - 1	:	.1930	0.	0.		.1481	.0091	0.		
1 - 2	:	.2050	0.	0.	:	.1122	.1082	.0001		
$\frac{1}{2} - \frac{1}{3}$		.1204	0.	0.	:	.1083	.0305	.0365		
3 - 4	:	.1572	0.	0.	:	.0739	.0200	.0315		
4 - 5	:	.1002	.0217	0.	:	.0639	.0102	.0219		
5 - 6	:	.0743	.0090	0.	:	.0425	.0045	.0582		
6 - 7	:	.0874	.0152	0.	:	.0481	.0253	.0754		

Table 4. Energy deposition components (GeV) for 120 GeV proton.

Component		Tungst	en	:	Сор	per
		MARS10	FLUKA86	:	MARS10	FLUKA86
Ionization	:	.93	1.05	:	.33	.39
EMS	:	9.30	8.35	:	.90	.73
Nucl.excitation	1:	. 55	.94	:	.11	.25
$\mathbf{E}$ < $\mathbf{E}\mathbf{th}$	:	_	.22	:	_	.07
Leakage	:	109.0	109.5	:	118.	117.8

<sup>\*)</sup> The nuclear excitation component includes total excitation energy in FLUKA86 and only energy deposition from particles after de-excitation stage in MARS10.

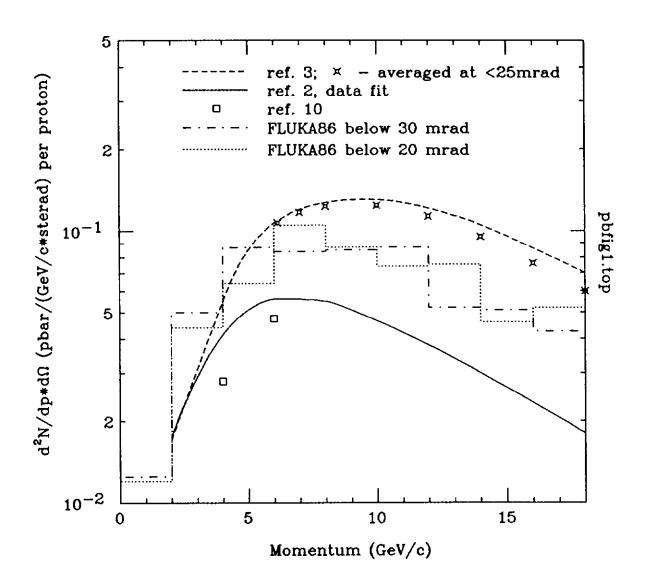


Fig.1. The pbar momentum distribution at zero angle for 70 GeV/c proton-tungsten interaction as predicted by various authors.

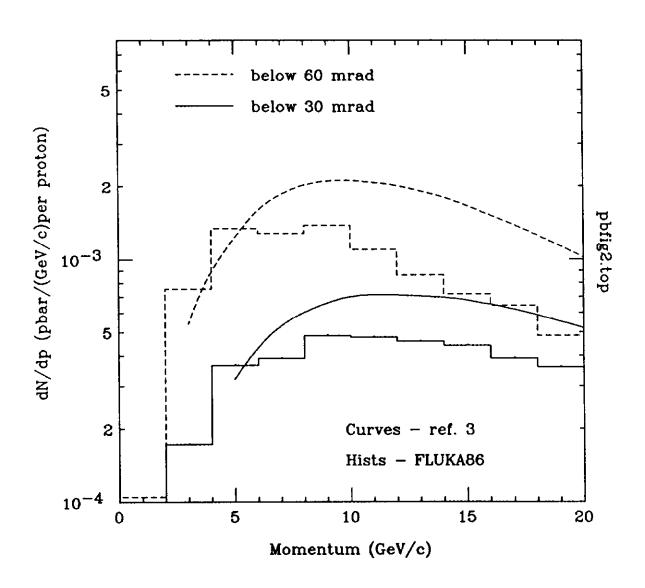


Fig.2. Integrated over two angular bins antiproton momentum distribution for 120 GeV/c pW interactions from ref.[3] and calculated in the present paper.

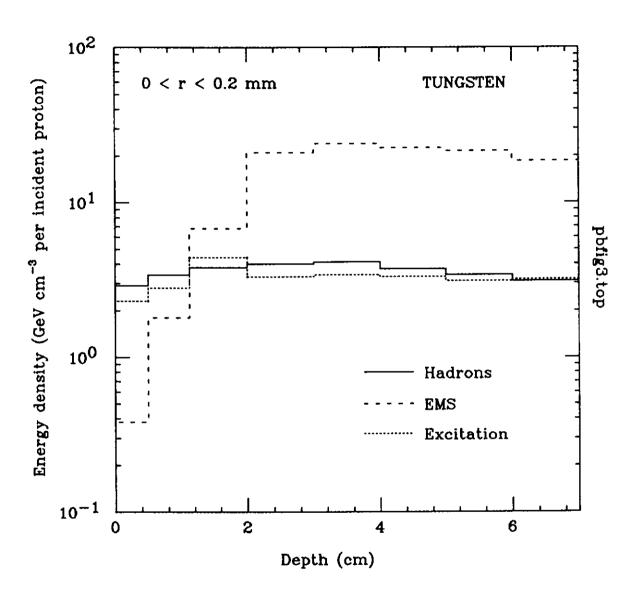


Fig.3. Longitudinal distibutions of energy density components around the beam axis in the Fermilab Pbar Source tungsten target at 120 Gev/c as calculated with MARS10.

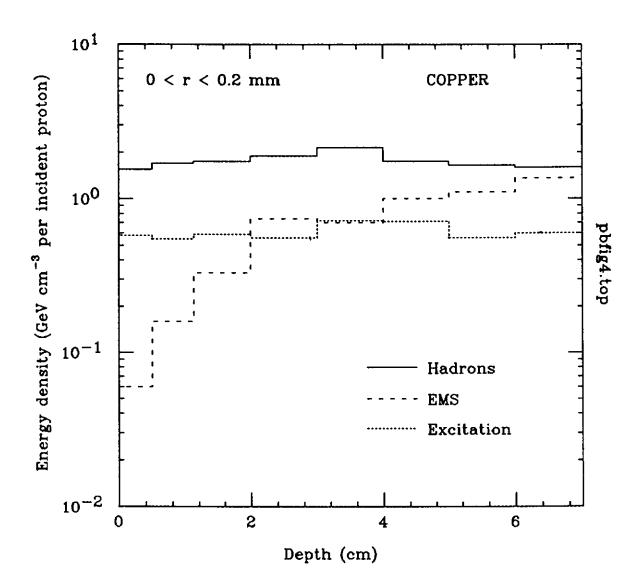


Fig.4. The same, as in fig.3, for copper target.

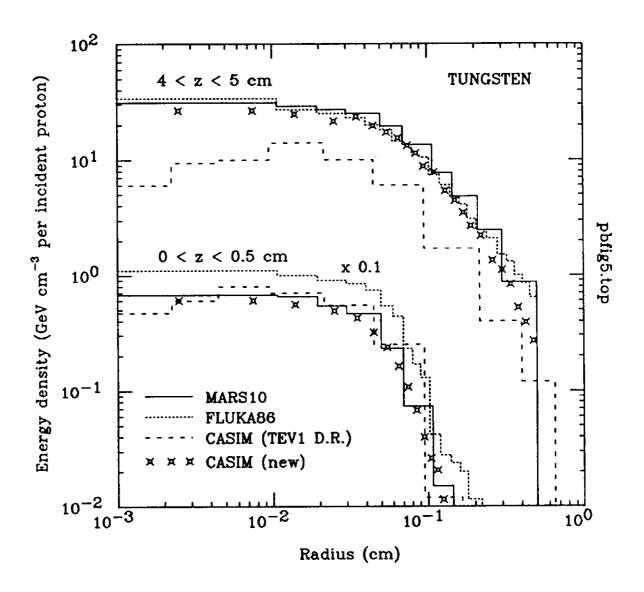


Fig.5. Radial distibutions of energy density at two depths of the Pbar tungsten target as calculated with three programs. The upstream results are multiplied by  $\emptyset.1$ .

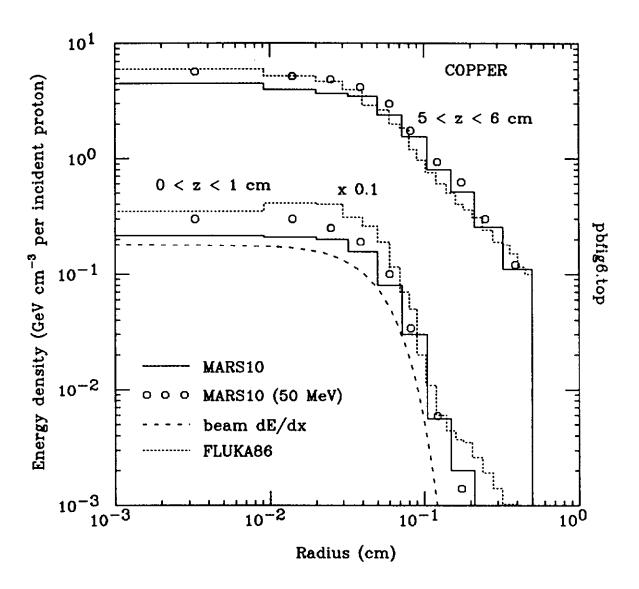


Fig.8. The same, as in fig.5, for copper target calculated with the MARSIO and FLUKA88 programs. For the small depth also shown are the MARSIO data obtained with the increased as in FLUKA88 threshold energies and ionization losses of primary protons.

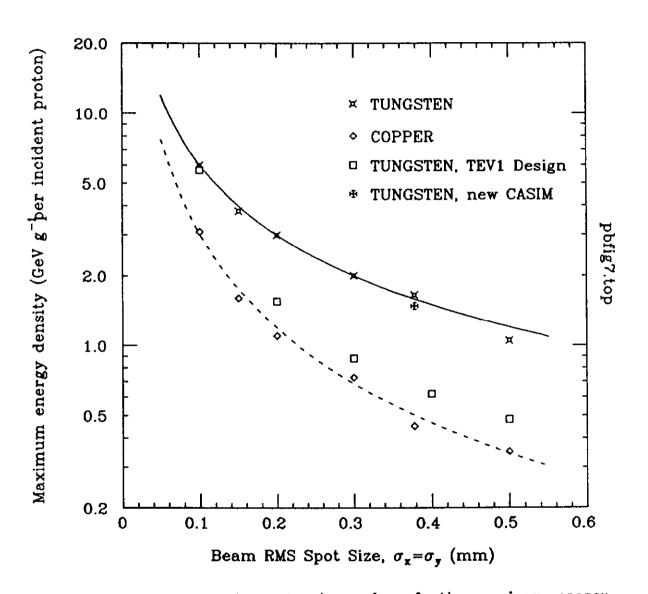


Fig.7. Dependence on the proton beam size of the maximum energy density in the Phar tungsten and copper targets as calculated with MARSIO. Curves are the fits described in the text. Also shown for tungsten are the Design Report data [4] and the new CASIM result.

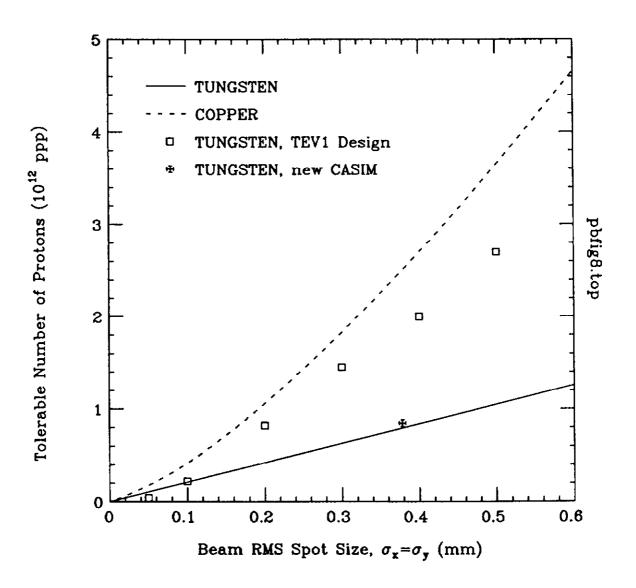


Fig.8. Tolerable number of protons on the Pbar tungsten and copper targets at 120 GeV/c as function of the beam spot size. Also shown are tungsten data from ref.[4] and the new CASIM result.